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The geochemical signatures of variable gas venting at gas hydrate sites

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Abstract

Diverse evidence suggests that gas-venting rates at sites of hydrate crystallization are variable in space and time, but the magnitude of these variations has been difficult to quantify. The hydrate crystallization model of Chen and Cathles [J. Geophys. Res. (Solid Earth) 108 (2003)] is used here to analyze 10 years of vent gas chemistry measurements at the Bush Hill hydrate mound and gas-venting site, Green Canyon 185, offshore Louisiana, Gulf of Mexico. The analysis suggests that, at any instant of time, gas vents at variable rates in different gas channels at the same site, and that the compositional differences in these vent gases are nearly as large as can be produced by hydrate crystallization. Almost two orders of magnitude differences in venting rate between individual gas channelways are suggested. Changes in the average vent gas composition over the last 10 years suggest the average venting rate varied by a factor of ~ 2 or more over a few years. The average $C_3 + C_4$ composition of Bush Hill hydrates are leaner than could be crystallized from vent gases sampled over the last decade, indicating that the venting gas flux was slower in the past by a factor of ~ 2 . This is compatible with geologic generalizations that venting evolves from fast (mud volcano), to intermediate (hydrate crystallization), to slow (carbonate precipitation) if venting organized into more discrete vents with time.

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Keywords: Chemical signature; Variable gas venting; Vent and hydrate gas composition; Bush Hill; Gulf of Mexico

1. Introduction

Gas hydrate is an ice-like crystalline mineral in which a rigid cage of water molecules encloses hydrocarbon and non-hydrocarbon gas molecules (Sloan, 1998). Natural gas hydrates occur worldwide in Polar Regions and in subthermocline oceanic environments, especially in areas of onshore and offshore permafrost and in sediments on continental margin slopes (Kvenvolden, 1998). The Gulf of Mexico is a classic area of gas hydrate occurrence. Hydrates have been sampled at more than 50 sites where the water depth exceeds ~440 m (Booth, Rowe, & Fischer, 1996; Kennicutt et al., 1985; Kvenvolden & Keith, 1995; MacDonald, Guinasso, Sassen, Brooks, Lee, & Scott, 1994; Sassen & MacDonald, 1994; Sassen et al., 2001a,b). Milkov & Sassen (2001) estimate that there is $10-14 \times 10^{12}$ m³ of gas entrapped in hydrate in the northwestern Gulf of

Mexico, and suggest that $\sim 80\%$ occur near faults at the margins of salt withdrawal minibasins where thermogenic gas has vented.

Many of the thermogenic, fault-related hydrate sites in the Gulf of Mexico are also sites of active gas and oil venting. MacDonald et al. (1993) documented 60 active oil seeps in the northern Gulf by analysis of satellite remote sensing data, but recent, more comprehensive data indicate that this estimate is too low by an order of magnitude (MacDonald et al., 2000; 2002). Manned submersible observations and seafloor mapping has documented a range of venting rates with characteristic physical, chemical and biological features (Roberts, 2001; Roberts & Carney, 1997). The most rapidly erupting vents disgorge massive quantities of sediment and hydrocarbon rich fluids that produce mud volcanoes and mudflows. Gas venting at intermediate rates supports viable and widespread chemo-synthetic communities, crystallization of surface and near-surface gas hydrates, and spatially variable small-scale mud volcanoes. Slow gas seepage is associated with mounded or

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chimney-like authigenic carbonates or other less impressive 113 forms of seafloor lithification, such as hardgrounds, slabs, 114 and nodules, and exotic minerals like barite in localized 115 areas. Although not necessarily always the case, these 116 systems have evolved from fast to slow venting in at least 117 some cases. In the Green Canvon Block 338 area, for 118 119 example, barite chimneys characteristic of slow gas venting are establishing themselves on Pleistocene mud flows 120 121 (Roberts & Carney, 1997).

122 The Bush Hill hydrate mound and gas vent is the best-123 studied intermediate vent site in the Gulf of Mexico and 124 perhaps the world. It is located where antithetic faults to a 125 major growth fault system intersect the sea floor in Green 126 Canyon Block 185 (Brooks, Kennicutt, Fay, MacDonald, & 127 Sassen, 1984, Brooks et al., 1986). The main fault system 128 contains the Jolliet oil and gas reservoirs (Cook & D'Onfro, 129 1991; Sassen et al., 2001a). Present-day gas venting at Bush 130 Hill produces a plume that is dramatically visible on echo 131 sounder records. Gas bubbles 2-3 cm in diameter breach 132 the sea surface over the site, leaving oil slicks as they 133 dissipate (Sassen et al., 2001a). Bush Hill hosts abundant 134 chemosynthetic communities, authigenic carbonates, and 135 bacterial mats. Hydrates are exposed on the surface of a 136 mound 800 m in diameter. Numerous sites of gas venting 137 are situated on this mound. Gases and hydrates at these vent 138 sites have been sampled by submersibles at least five times 139 in the last decade (see Table 1). A separate mud vent site lies 140 to the side of the Bush Hill hydrate mound. 141

The composition of the vent gas at Bush Hill is 169 consistent with stripping of C₂₊ hydrate-forming gases 170 from Jolliet reservoir gas by hydrate crystallization 171 (Sassen et al., 2001a). Since there is no isotopic 172 fractionation during hydrate crystallization, the isotopic 173 similarity of the Jolliet reservoir gases, the vent gases, 174 and the Bush Hill hydrates is very strong evidence that 175 all the gasses are from a common source (Roberts, 2001; 176 Sassen et al., 2001a). Based on this common source, a 177 kinetic model of gas venting and hydrate crystallization 178 at the Bush Hill was constructed by Chen and Cathles 179 (2003). Analysis of the Bush Hill vent with this model 180 showed that, on average over the 10 000 year history of 181 hydrate accumulation, $\sim 9\%$ of the vent gas is crystal-182 lized as hydrate in the subsurface and the compositions 183 of vent gases and hydrates are controlled almost entirely 184 185 by the variation in the rate of gas venting. The purpose 186 of this paper is to examine the changes in vent gas 187 composition that have been observed over the last 10 years at Bush Hill in more detail, and to suggest how 188 189 these variations could be related to vent evolution. The 190 changes in vent chemistry suggest a dynamic, constantly 191 changing nexus of subsurface gas pathways that are 192 progressively organized so that venting occurs at fewer 193 locations of smaller total area as the overall venting rate 194 gradually wanes. This progressive vent consolidation is 195 an aspect of gas vent evolution that to our knowledge 196 has not been previously specifically discussed. 197

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143	Table	1

144	Observed	composition	of Bush	Hill	vent	and	hydrate	gase
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No.	Year	<i>T</i> (°C)	C ₁	C ₂	C ₃	i-C ₄	n-C ₄	i-C ₅	n-C ₅	$C_3 + C_4$
V-a	1993	7.35	88.0	8.0	2.1	0.3	1.2	0.4	< 0.1	0.036
V-b	1993	7.35	88.0	7.5	2.2	0.5	1.1	0.6	< 0.1	0.038
V-1	1995	9.10	93.2	4.3	1.5	0.3	0.6	0.3	< 0.1	0.024
V-2	1995	9.10	93.5	4.3	1.4	0.2	0.4	0.2	< 0.1	0.020
V-3	1995	9.10	94.7	3.9	0.7	0.1	0.5	0.2	< 0.1	0.013
V-4	1995	9.10	94.6	3.8	0.7	0.1	0.5	0.3	< 0.1	0.013
V-5	1995	9.10	91.1	4.8	1.8	0.4	1.2	0.8	< 0.1	0.034
V-6	1997	7.00	90.4	4.5	3.7	0.6	0.6	0.2	< 0.1	0.049
V-7	1997	7.00	95.9	2.4	1.2	< 0.1	0.3	0.2	< 0.1	0.015
V-8	1998	7.00	93.4	4.1	1.5	0.3	0.5	0.3	< 0.1	0.023
V-9	2000	6.50	92.7	4.6	1.7	0.3	0.6	0.2	< 0.1	0.026
Mean		7.00	92.32	4.75	1.68	0.31	0.68	0.34	< 0.1	0.02645
	H-1		83.1	7.6	8.10	0.9	0.2	0.0	0	0.092
	H-2		71.7	10.6	12.6	2.6	1.7	0.8	0	0.169
	H-3		80.2	9.40	7.30	1.6	1.2	0.3	0	0.101
	H-4		72.1	12.4	11.4	2.3	1.6	0.3	0 0	0.153
	H-5		85.7	6.30	6.10	1.1	0.8	0.0	0	0.080
	H-6		71.8	3.40	18.8	5.7	0.3	ND	ND	0.248
	H-7		73.9	4.90	16.3	4.6	0.2	ND	ND	0.211
	H-8		72.1	10.5	12.4	2.5	1.7	< 0.1	0.7	0.166
	Mean		76.33	8.1375	11.625	2.6625	0.9625	0.175	0.0875	0.1525

167 223 MacDonald et al., 1994; Sassen and MacDonald, 1994, 1997; Sassen et al., 1998, 1999a,b. All temperature data and vent gas sample-V-9 is unpublished 168 data from Sassen, R. 224

225 2. Chemical analysis of Bush Hill venting

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Table 2

2.1. Kinetic model of gas venting and hydrate crystallization 227 in Bush Hill 228

The kinetic model for hydrate crystallization from a gas 230 stream is fully presented in Chen and Cathles (2003). The most important aspects of this model are reviewed here. The 232 model addresses the rate at which gas is crystallized from a 233 gas stream, (DM/Dt), in a coordinate system moving with 234 the gas: 235

$$\frac{236}{237} \qquad \frac{DM}{Dt} = -k(X_{3+4}^{\nu\text{-gas}} - X_{3+4}^{\text{equ}})\exp\left(\frac{E^*}{R}\left(\frac{1}{T^*} - \frac{1}{T}\right)\right). \tag{1}$$

239 Here k (kg/m³ a) is the product of the kinetic rate constant 240 and the actively crystallizing hydrate surface area per unit 241 volume, $X_{3+4}^{v-\text{gas}}$ is the C₃ + C₄ mass fraction composition of 242 the vent gas, X_{3+4}^{equ} is the C₃ + C₄ mass fraction composition 243 of the gas that would be in equilibrium with hydrate at the 244 pressure and temperature selected, E^* is the activation 245 energy of the reaction, R is the gas constant, T^* is an 246 arbitrary reference temperature which was taken to equal 247 273.15 K, and T is the temperature at the location of the gas 248 packet in K.

249 Eq. (1) applies at depths below the sea floor where gas 250 hydrate is stable (e.g. $X_{3+4}^{\nu-\text{gas}} - X_{3+4}^{\text{equ}} > 0$). It is solved by 251 propagator methods. First we use the methods of Sloan 252 (1998) to calculate the subsurface depth interval over which 253 hydrate can crystallize, which we will refer to subsequently 254 as the potential hydrate stability zone (HSZ). This 255 calculation is made, assuming pore fluids have seawater 256 salinity, by determining the depth at which gas with the 257 composition of a selected Jolliet reservoir gas, $X_{3+4}^{J-\text{gas}}$, begins 258 to crystallize hydrate. The potential HSZ extends to the 259 surface because we assume that venting can be very fast and 260 because gas discharge is observed at Bush Hill. The HSZ is 261 then divided into a number of depth intervals. The gas is 262 moved, unchanged in composition to the middle of the first 263 depth interval in the HSZ, X_{3+4}^{equ} is calculated for the pressure 264 and temperature there, and the rate of hydrate crystallization 265 is computed from Eq. (1). The fractional crystallization of 266 gas that occurs in the time the gas takes to transit the layer 267 (at a mass flux q (kg/m² a)) is then computed, and the 268 composition of vent gas is adjusted to account for hydrate 269 crystallization. The composition of the hydrate crystallized 270

is a function of P, T and $X_{3+4}^{\nu-\text{gas}}$ (Chen & Cathles, 2003; 281 Sloan, 1998). The gas is then introduced to the next layer 282 and the calculations repeated. 283

2.2. The causes of vent gas compositional changes at Bush Hill

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288 In the broadest interpretive context, the analysis of vent 289 gas compositions predicted here shows that (of the variables 290 considered) only variations in venting rate could account for 291 the observed variations in Bush Hill vent gas composition. 292 Source gas compositions cannot account for the chemistry 293 of the vent gases because the $C_3 + C_4$ range of Jolliet 294 reservoir gas does not span the observed range in vent gas 295 composition. With venting-rate variation, a feed gas with 296 the average Jolliet reservoir gas composition can account for 297 all but the wettest (most enriched in $C_3 + C_4$) observed vent 298 gases at Bush Hill. For these few highly enriched vent gases 299 a source gas slightly wetter than the average Jolliet reservoir 300 gas is required. We constrain these conclusions by 301 calculating the range of vent gas compositions that could 302 result by varying all the controlling parameters over their 303 full possible range using the methods of Chen and Cathles 304 (2003).

305 Four observed parameters influence vent gas composition: source gas composition $(X_{3+4}^{J-\text{gas}})$, seafloor temperature 306 307 $(T_{\rm sf})$, subsurface temperature gradient (G), and gas-venting 308 rate (q). The ranges of these parameters are shown in Tables 309 1 and 2. The mass fraction of $C_3 + C_4$ in Jolliet reservoir gas 310 (the assumed source gas) ranges from 0.031 to 0.06. 311 Seafloor temperatures of 6-7 °C are normal for 540 m 312 water depth (Walker, Huh, & Rouse, 1993) and tempera-313 tures recorded at the seafloor at 540 m in GC 185 range 314 between 6 and 11 °C with a mean of 7 °C (MacDonald et al., 315 1994; Sassen & MacDonald, 1994). Loop current eddies 316 spinning off from the Gulf Stream can have cores that reach 317 14 °C at a water depth of 500 m (MacDonald et al., 1994; 318 MacDonald, Buthman, Sager, Peccini, & Guinasso, 2000; 319 Roberts, 2001; Roberts & Carney, 1997). Seafloor tempera-320 ture ranges from 6 to 14 °C at Bush Hill. Eighteen Jolliet 321 reservoir temperature measurements updated in Bascle, 322 Nixon, and Ross (2001) and 19 subsurface measurements 323 define the geothermal gradient in Bush Hill area at 20.1 °C/ 324 km (Chen & Cathles, 2003). The geothermal gradient can, 325 however, be affected by rapid vertical fluid flow (Roberts, 326

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$X_{3+4}^{J-\mathrm{gas}}$	0.031-0.06	0.031-0.06	0.031-0.06	0.04675	0.04675	0.04675
$T_{\rm sf}$ (°C)	5-15	7	7	7	5-15	5-15
G (°C/km)	20	15-50	20	15-50	20	15 - 50
$q (\text{kg/m}^2 \text{ a})$	1.84	1.84	0.55-55.26	0.55-55.26	0.55-55.26	1.84

279 335 $X_{3+4}^{J_{283}}$ is the measured range of Jolliet reservoir gases (Sassen et al., 2001a); T_{sf} is the seafloor temperature in °C; G is the subsurface geothermal gradient in 280 $^{\circ}$ C/km; q is gas mass flux in kg/m² a. The numbers in the first row correspond to the x-axis labels of the vertical line bars in Fig. 1. 336

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2001; Roberts & Carney, 1997). Observation at other sites 337 show that rapid pulses in mud discharge can increase the 338 surface temperature by >10 °C (MacDonald et al., 2000; 339 Roberts, 2001). A range of the subsurface geothermal 340 341 gradient from 15 to 50 °C/km should span all plausible 342 subsurface geothermal gradients at Bush Hill. Venting rates that impact hydrate composition vary from 0.55 to 343 344 55 kg/m² a as we will show below.

Fig. 1 shows the consequences for vent gas composition of varying the six possible pairs of these four parameters $(X_{3+4}^{J\text{-gas}}, T_{sf}, G, \text{ and } q \text{ as defined in Table 2})$ that control vent gas composition. The observed ranges in vent gas and hydrate composition are shown as gray bands. The black bars and circles indicate the range in reservoir gas and the mean reservoir gas composition, respectively.

The first thing to notice in Fig. 1 is that the range in Jolliet reservoir gas composition (lower black bar) covers at most half of the range in Bush Hill vent gas composition that has been measured over the last 10 years, and that the gases in hydrate are all heavier than the reservoir gases (upper black bar). The heavier gas components preferentially fractionate into hydrate as it crystallizes. The reservoir gases can thus 393 be converted to the compositions of the vent gases by 394 crystallizing hydrate (Sassen et al., 2001a). If the venting is 395 very rapid, little gas will be crystallized, and the vent gas 396 will have very nearly the same composition as the reservoir 397 gas. The heaviest vent gas (top of gray band in the lower 398 plot) is just slightly heavier (more enriched in $C_3 + C_4$) than 399 the mean Jolliet reservoir gas (black dot on lower figure). 400 Variation in venting rate (and/or other factors) could thus 401 account for nearly the full range in observed vent gas 402 compositions, but some heavier-than-average reservoir gas 403 must contribute to some vent pathways to account for the 404 405 vent gases heavier than the mean Jolliet reservoir gas.

406 The vertical lines in Fig. 1 show the vent gas and hydrate compositions predicted by our model where the indicated 407 408 pairs of variables are varied across their permissible ranges 409 as shown in Table 1. Only the vertical lines that include the 410 gas-venting rate, q, as a variable cross the full observed 411 range of either hydrate or vent gas composition. Although to 412 span the full range in vent gas composition, the source gas 413 composition must be varied slightly, the most important 414

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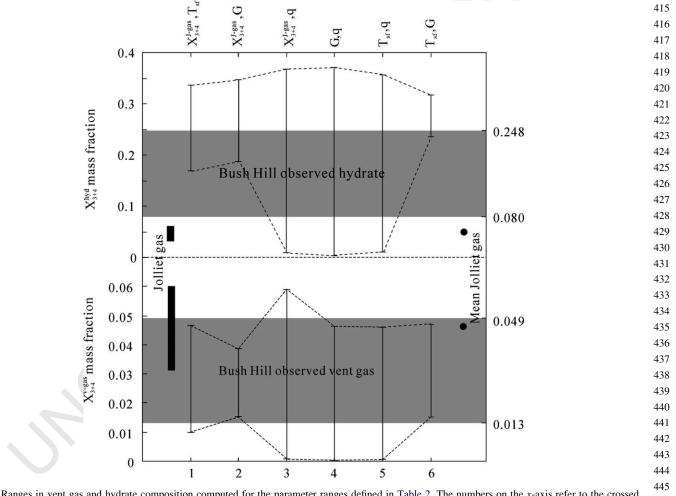


Fig. 1. Ranges in vent gas and hydrate composition computed for the parameter ranges defined in Table 2. The numbers on the *x*-axis refer to the crossed parameters in Table 2. The crossed parameters that are varied across their permissible range are shown above the top axis. Gray shading shows the range of observed vent and hydrate gas compositions. Black dots in both diagrams show the mean Jolliet reservoir gas composition. Black bars in both diagrams show the composition range of Jolliet reservoir gases.

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conclusions to draw from Fig. 1 is that, of the factors
considered, variation in venting rate is the most important
factor controlling vent gas and hydrate composition. This
factor alone can account for nearly the entire range of vent
gas and hydrate composition.

2.3. The compositional span of vent rate variations

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Fig. 2 plots (solid line) the predicted compositions of vent gas and hydrate crystallized near the seafloor as a function

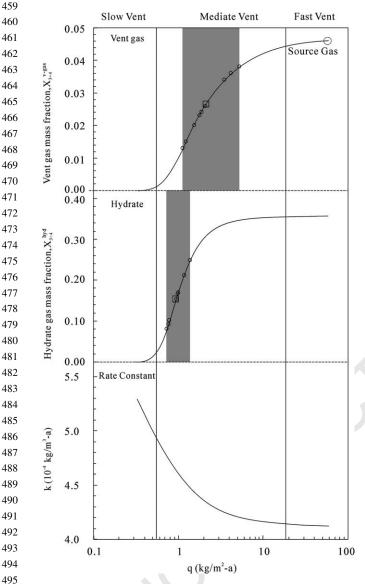


Fig. 2. Vent gas and hydrate compositions calculated as a function of gas vent 496 rate for $T_{\rm sf} = 7$ °C, G = 20 °C/km, and a source gas composition equal to the 497 mean Jolliet reservoir gas. The boundaries of fast, intermediate and slow 498 venting are indicated by vertical lines. Gray shading indicates the compositional range of sampled vent gases $(X_{3+4}^{v-\text{gas}} = 0.013 - 0.049 \text{ with})$ 499 corresponding computed gas mass flux q = 1.13-5.21 kg/m² a) and hydrate 500 gases $(X_{3+4}^{hyd} = 0.08 - 0.248$ with corresponding computed gas mass flux 501 $q = 0.73 - 1.38 \text{ kg/m}^2 \text{ a}$) at Bush Hill. Open cycles are the sampled vent and 502 hydrate gas compositions and squares are their average values. The bottom 503 figure shows how the rate constant must be adjusted if a hydrate mound 504 containing an average 2 vol% hydrate is to accumulate in 10 000 years.

of gas vent rate, q. Measured vent gas and hydrate 505 composition are indicated by symbols, and the range of 506 observed compositional data is indicated by gray bands. The 507 calculations assume a surface temperature of 7 °C, a 508 temperature gradient of 20 °C/km, and a pore salinity of 509 3.54%. An additional constraint used in constructing this 510 figure is that after 10 000 years of venting, the average 511 hydrate accumulation equals 2 wt% of the sediment volume 512 as required by observations and geological constraints at 513 Bush Hill (see, Chen & Cathles, 2003). All thermodynamic 514 calculations are made using the CSMHYD program (Sloan, 515 1998) as discussed in Chen and Cathles (2003). 516

At very slow model venting rates, the heavy gas 517 components are all crystallized as hydrate below the sea 518 floor, and the gas wetness (as measured by the $C_3 + C_4$ mass 519 fraction) is zero in both the vent gas and hydrate. At model 520 gas mass fluxes greater than $\sim 0.5 \text{ kg/m}^2$ a, vent gas and 521 surface hydrates contain a sufficient mass fraction of 522 $C_3 + C_4$ hydrocarbons that hydrate can crystallize. The 523 compositions of vent gas and hydrate depend strongly on 524 venting rate up to venting rates of $\sim 20 \text{ kg/m}^2$ a. At very fast 525 model venting rates, the vent gas approaches the source gas 526 composition (average Jolliet reservoir gas in this case), and 527 the hydrates crystallized near the surface are as heavy as 528 thermodynamics and the source gas wetness allow. 529

The constraint that the average hydrate accumulation 530 equal 2 wt% after 10 000 years of venting was imposed by 531 varying the kinetic rate constant, k. The bottom panel in 532 Fig. 2 shows how k had to be varied to meet this constraint. 533 The range in k required is quite small (from 4.1 to 534 5.3×10^{-4} kg/m³ a). This relationship illustrates that, as 535 discussed in Chen and Cathles (2003), the kinetic rate 536 constant is not a particularly important variable once it is 537 empirically calibrated. The rate of hydrate crystallization 538 does not depend strongly on venting rate until the venting 539 rate becomes so weak that near-surface hydrate crystal-540 lization is strongly impaired. 541

Finally, Fig. 2 shows that the mass fractions of $C_3 + C_4$ 542 hydrocarbons in hydrates at Bush Hill suggest a slower 543 venting rate than do vent gases collected over the last 10 544 years (e.g. the gray band enclosing the hydrate compositions 545 lies to the left of the gray band enclosing the vent gas 546 compositions). If the hydrates are older than 10 years this 547 suggests the venting rate has been increasing with time. At 548 face value, this contradicts intuition that a gas blow out from 549 the overpressured zones in a basin should be rapid at first 550 and then decline, and the geological inferences of Roberts 551 and his colleagues (Roberts, 2001; Roberts & Carney, 1997) 552 that the gas-venting rate decreases with time as a vent 553 matures from the mud volcano stage through hydrate/ 554 chemosynthetic to carbonate stage. These stages are roughly 555 demarcated by the vertical lines in Fig. 2. This apparent 556 discrepancy can be resolved if the area of venting collapses 557 with time, as discussed below. 558

Table 3 summarizes additional features of the three zones559(slow, medium, and fast venting) in Fig. 2. For fast venting,560

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Table 3 Boundaries of fast, ir	ntermediate and slow venting	classes as defined by observed	and computed parameter	rs	
Vent type	$X_{3+4}^{v-\mathrm{gas}}$	X ^{hyd} ₃₊₄	ΣF (%)	F_{10} (%)	$q (\text{kg/m}^2 \text{a})$
Fast	>0.044	>0.355	< 0.9	0.1	>18.42
Intermediate Slow	0.001 - 0.044 < 0.001	0.022 - 0.355 < 0.022	0.9-23 > 23	0.1–2.8 No hydrates	0.55 - 18.42 < 0.55

 $\begin{array}{ll} 568 & X_{3+4}^{y\text{-gas}} \text{ is the mass fraction } C_3 + C_4 \text{ in the vent gas}, X_{3+4}^{hyd} \text{ is the mass fraction } C_3 + C_4 \text{ in the hydrate crystallised at or near the surface, } \Sigma F \text{ is the total fraction} & 624 \\ 569 & \text{of venting gas crystallized as hydrate}, F_{10} \text{ is the fraction of venting gas crystallized in the uppermost } 10\% \text{ of the hydrate stably zone} (surface to ~60 \text{ meters} \\ 570 & \text{deep}), \text{ and } q \text{ is the gas mass flux (venting rate) in kg/m^2 a. The venting rate is calculated assuming there is 2 vol% hydrate in a numerical Bush Hill hydrate \\ 571 & \text{mound that is } 600 \text{ m diameter at the surface and extends to depth as calculated in Chen and Cathles (2003).} \\ \end{array}$

574 the venting gas and surface hydrates are rich in $C_3 + C_4$ 575 hydrocarbons, a small fraction of the gas crystallizes as 576 hydrates in the subsurface (< 0.9%), and a proportional 577 amount of hydrate crystallizes in the uppermost 10% of the 578 HSZ (0.1% of the gas flux). At intermediate venting rates, 579 the vent gas and hydrate compositions depend strongly on 580 venting rate, up to 23% of the venting gas crystallizes as 581 hydrate, and a proportional amount still crystallizes in the 582 shallowest 10% of the HSZ. At slow venting rates over 23% 583 of the gas stream crystallizes as hydrate in the subsurface, 584 but the hydrate crystallization is all at substantial depth. No 585 hydrate crystallizes in the shallowest 10% of the HSZ. Note 586 we arbitrarily define the HSZ here with reference to the 587 source gas composition. It extends from the depth at which 588 hydrate is first stable for this source gas to the surface 589 regardless of venting rate. 590

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Fig. 3 shows how the model vent gas and model hydrate compositions change with depth as a function of venting

rate. The solid lines indicate the model venting rates that match the observed compositional data in Fig. 2. This figure again emphasizes that the hydrate samples suggest a slower venting rate than do the vent gas samples collected in the last 10 years.

2.4. Interpretation of a decade of Bush Hill vent gas chemistry

Vent gas and hydrate samples have been collected during multiple submersible visits to Bush Hill since 1993. Table 1 lists all available data. The hydrate samples are of indeterminate age and show signs of (minor) bacterial oxidation (Sassen et al., 1999a). The mound appears to have accumulated over several thousand years. Carbonates in the mound have been dated at 1.4-3.2 ka (Roberts & Aharon, 1994). The C₃ + C₄ mass fraction of vent gases are plotted as a function of sampling time in

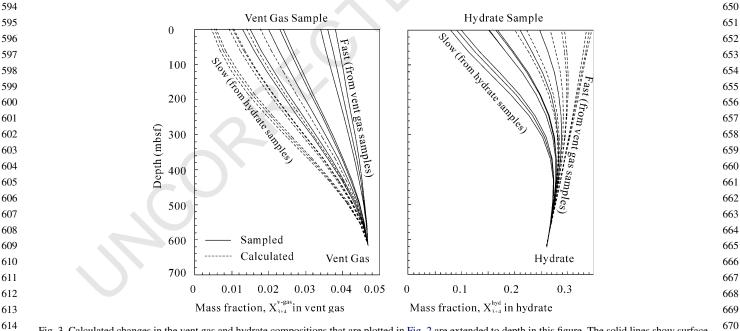


Fig. 3. Calculated changes in the vent gas and hydrate compositions that are plotted in Fig. 2 are extended to depth in this figure. The solid lines show surface data extrapolated to depth using the model. The dashed lines are for reference. They show the extrapolated lines from the other figure (e.g. superimposed $X_{3+4}^{\nu,\text{gas}}(z)$ on the $X_{3+4}^{hyd}(z)$ plot and visa versa).

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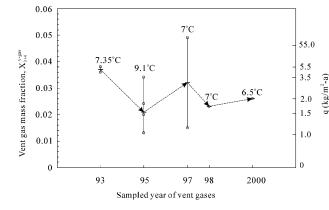


Fig. 4. Bush Hill vent gas compositions plotted against the year sampled. Seafloor temperatures at the time of sampling are shown. The right axis shows the gas mass fluxes suggested by the vent gas compositions.

Fig. 4. The variation in vent gas composition at any particular sampling time is greater than the variation between samplings. The mean mass fraction $C_3 + C_4$ composition changes by a factor of ~2 between submersible samplings.

There are very little data, but the implications of these data in terms of the modeling analysis presented are profound. If the composition differences between different bubble streams sampled at the same site on the same submersible visit are the result of variations in the mass flux of the individual bubble streams, an almost two order of magnitude variation in mass flux is required (right hand vertical axis in Fig. 4). The data suggest that at any instant of time gas may be venting from depth to the surface through a number of independent channelways at very different rates. The gas venting from channelways where the venting rate is slow is depleted in $C_3 + C_4$, while the channelways in which the venting rate is fast have $C_3 + C_4$ mass fractions approaching those in the source reservoir gas. The changes in vent gas composition between submersible visits suggest that the venting rate (gas mass flux) in individual channels is also changing with time.

Table 4 provides the details of these calculations. It shows the total fraction of gas, ΣF , that was crystallizing as hydrate from each bubble stream when it was measured, the composition of the surface hydrate crystallizing from the gas, and the gas mass flux calculated for the bubble stream for the seafloor temperature measured at the time of sampling (as listed in Table 1) or a sea floor temperature of 7°C if none was measured.

2.5. Synthesis

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Fig. 5 suggests a synthesis of the analysis presented above and the geologically indicated evolution of vents suggested by seafloor observations summarized in Roberts (2001) and Roberts and Carney (1997). This figure depicts

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No. Year	$X_{3+4}^{v\text{-gas}}$	ΣF	X_{3+4}^{hyd}	$q (\text{kg/m}^2 \text{ a})$		
				q_a	q_b	
V-a	1993	0.036	0.041	0.342	4.19	3.83
V-b	1993	0.038	0.033	0.346	5.21	4.74
V-1	1995	0.024	0.089	0.295	1.87	0.99
V-2	1995	0.020	0.106	0.268	1.54	0.81
V-3	1995	0.013	0.141	0.202	1.13	0.59
V-4	1995	0.013	0.141	0.202	1.13	0.59
V-5	1995	0.034	0.049	0.337	3.50	1.89
V-6	1997	0.049	None	None	None	None
V-7	1997	0.015	0.130	0.224	1.23	1.23
V-8	1998	0.023	0.093	0.289	1.78	1.78
V-91	2000	0.026	0.080	0.306	2.07	2.32
Mean	1993-2000	0.0264	0.078	0.309	2.12	2.12

 ΣF is the total fraction of venting gas crystallized as hydrate, X_{3+4}^{hyd} is the $C_3 + C_4$ mass fraction of hydrate crystallized at the seafloor. Gas mass flux, q_a , is calculated for a mean seafloor temperature of 7 °C; gas mass flux, q_b , is calculated for the sea floor temperatures measured at the time the samples were collected (see Table 1).

751 the early, mud volcano stage of venting as having the 752 753 largest total gas venting Q, but shows this venting spread over a broad area (Fig. 5a). Because the venting is spread 754 755 over a broad area, the gas mass flux, q = Q/A, is low. 756 Hydrates crystallize from the base of the HSZ to the 757 surface, but because rapid mud flow may warm the 758 subsurface (and even the sea floor) the depth interval of 759 hydrate crystallization may be slightly reduced as shown 760 by the dashed line in the figure. Fig. 5b depicts the gas 761 venting as focused to a much smaller area at the hydrate/ 762 chemosynthetic stage of venting. The total gas discharge 763 rate, Q, is smaller in the figure but the gas flux, q, is 764 larger because the area of discharge is reduced more than 765 the total venting Q is reduced. Finally, when the venting 766 rate drops below that required to sustain hydrate crystal-767 lization at the surface venting occurs only in a very few 768 widely scattered locations and the vent enters its 769 carbonate evolutionary stage (Fig. 5c). In this stage, the 770 carbon tied up in hydrates or dissolved in pore water may 771 be slowly converted to carbonate, producing carbonate 772 mounds and hardgrounds.

773 If this progressive channel organization occurred at Bush 774 Hill, it could explain the observed relationship between 775 hydrate and vent chemistry even if the overall venting rate 776 were decreasing with time. If the gas flux, q, controls the 777 vent and hydrate gas chemistry as suggested by our 778 modeling, the higher mass fraction $C_3 + C_4$ in vent gases 779 sampled in the last 10 years compared to the vent gas 780 compositions suggested by the (presumably older) hydrates 781 in the Bush Hill mound could reflect the progressive 782 organization of venting into channelways of smaller total 783 cross section and higher mass flux. 784

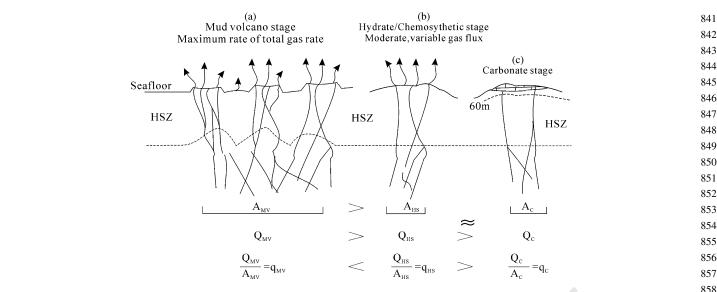
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Fig. 5. Cartoon showing the evolution of a typical gas vent. The total gas venting, Q, is most rapid at the mud volcano stage, but the venting occurs over a broad area, A, and the gas mass flux, q, is consequently low. Venting is progressively organized so the area of discharge is reduced, and q is increased. At the same time the total venting rate Q decreases. The three stages shown correspond to those geologically inferred by Roberts and co-workers as discussed in the text.

3. Discussion

The critical presumption in our analysis is that venting 808 rate and gas and hydrate chemistry are related. Bubble 809 streams fed by subsurface flow pathways through which 810 gas transit is fast discharge gases and crystallize hydrates 811 rich in $C_3 + C_4$. Bubble streams fed by gas flow pathways 812 with slow transit times discharge gas and crystallize 813 surface hydrates lean in $C_3 + C_4$ hydrocarbons. Kinetic 814 models of gas venting and hydrate crystallization display 815 these systematics. Although there are many uncertainties 816 in the models as discussed in Chen and Cathles (2003), 817 and the subsurface flow pathways may be complex (with 818 slow transit in some places and fast in others), it is hard 819 to imagine how kinetically controlled hydrate crystal-820 lization could avoid these systematics. We therefore 821 suggest that the very different chemistry of different 822 individual bubble streams at a site sampled during the 823 same submersible visit (see especially 1995 and 1997 in 824 Fig. 4) is due to the different rates at which the 825 discharging gas traversed the HSZ. 826

We cannot prove that the (HSZ) transits times of different 827 bubble streams are different, so this explanation for the 828 different chemistries is a suggestion. The suggestion can 829 however be tested in several ways. Fig. 3 shows that if 830 the vent gas chemistry is controlled by venting rate, the 831 variability in vent gas and hydrate chemistry should 832 decrease with depth to the variability of the source gas 833 (which is taken to be zero in this figure since the source gas 834 is the average Jolliet reservoir gas). This prediction could be 835 tested by drilling the Bush Hill hydrate mound. The 836 hypothesis that individual vents are variable in their venting 837 rate and that this directly affects the composition of the 838 vented gases could be tested by monitoring the venting rate 839 840 and composition of a single bubble stream. This experiment

has been attempted several times but for practical reasons (plugging of the instrument with hydrate, difficulty positioning and calibrating measuring devices, etc.) has proven to be very difficult. The importance of venting rate and its variations in time and space to our understanding of the hydrate system (as made clear by the models we have constructed) reinforces the importance of making these difficult measurements.

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Using the connection between venting rate and gas 870 and hydrate chemistry we further suggest that, because 871 virtually all the hydrate samples collected at Bush Hill 872 indicate (when analyzed by our models) the venting rates 873 were slower than at present when they crystallized, the 874 gas mass flux at Bush Hill is greater today than in the 875 past (e.g. the average transit time across the HSZ is less 876 today). We do not know that the hydrates sampled are 877 older than 10 years, but we know they are incompatible 878 with the present vent gas chemistry and cannot (as a 879 whole) have crystallized from it. Given that the hydrate 880 mound accumulated over thousands of years it is likely 881 the samples are older than 10 years. We also do not 882 know that some other process, such as bacterial attack, 883 may not have altered the hydrate chemistry. However, 884 judging from their isotopic alteration, the heavy com-885 ponents of the hydrates sampled at Bush Hill are not 886 altered whereas the C1 component is altered (somewhat) 887 in a fashion consistent with bacterial oxidation (Sassen 888 et al., 1999a). This suggests that bacteria attack the C_1 889 component of hydrate first, and that the Bush Hill 890 hydrates would be enriched in their $C_3 + C_4$ components 891 by whatever bacterial oxidation they have suffered. Our 892 chemical dilemma is that the hydrates are not rich 893 enough in $C_3 + C_4$ to be compatible with gases venting 894 over the last 10 years. Bacterial oxidation cannot help 895 resolve this dilemma because it shifts the hydrate 896

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compositions the wrong way. The simplest explanation 897 that we can think of is that the venting rates were slower 898 in the past than they have been over the last 10 years. 899 The venting rate at Bush Hill seems to be quite variable, 900 so this is perhaps no surprise. In addition to the data in Fig. 901 4, echo sounder images of the Bush Hill plume seem to be 902 quite variable. In August 2000, echo sounder images 903 showed the Bush Hill plume originated from a 600 m 904 interval at the seafloor centered on the Bush Hill hydrate 905 mound and nearly reached the sea surface. The image was 906 robust to echo sounder tuning and visually dramatic. In June 907 2002, the vent imaged by echo sounding was about one third 908 the size, less robust to tuning, and visually weaker. At both 909 visits, bubbles of gas were observed breaking the surface 910 leaving iridescent oil slicks. However, if there were no 911 systematic trends, venting variability should not show such 912 as systematic difference between present venting rates and 913 914 almost all hydrates sampled to date. It seems, therefore, that the current venting rate is unusually fast compared to that 915 which pertained over whatever prior period is spanned by 916 the hydrate samples. 917

On the face of it discharge rates more rapid today than in 918 the past contradicts both common sense and geological 919 920 inference, and thus warrants some discussion. There seems little question that the fluid venting at Bush Hill derives from 921 zones where fluid pressure is known to be above hydrostatic 922 starting at depths of \sim 6500 ft below the seafloor. Rupture 923 924 events, probably related to fault movements, tap these highpressure zones. It is hard to imagine that when this happens 925 the discharge would not be rapid at first and then decrease 926 with time. This is also suggested from the geomorphological, 927 geochemical and biological features summarized in Roberts 928 929 and Carney (1997). Admittedly, it is a jump to interpret the observed spectrum of venting rates to indicate a progression 930 from fast to slow venting in all cases. Bush Hill may not have 931 had a mud-volcano pre-cursor, and discharge rates could 932 have increased there as the result of recent fault movement or 933 seal rupture rather than because of a general progression from 934 fast to slow gas discharge. 935

However, there is another possibility that is suggested by 936 broad observations. Roberts and Carney (1997) describe 937 how the vents in slow discharge areas are weak, rare, and 938 scattered compared to sites venting at intermediate rates. It 939 seems that the number of vents tends to decrease as the 940 system evolves (assuming the intermediate sites evolve to 941 weak venting sites). If this is the case, it is possible that the 942 gas transit time could decrease even as the overall venting 943 rate decreases. We offer this as a possible way that the gas 944 mass flux rate at Bush Hill could have could have 945 systematically increased with time. 946

949 **4. Conclusion**

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This paper explores the relationship between the gas venting rate and the gas and hydrate chemistry in the Bush Hill vent site, Green Canyon Block 185, offshore Louisiana, 953 Gulf of Mexico using the kinetic model developed by Chen 954 and Cathles (2003). This model is applied to the Bush Hill 955 vent taking the widest plausible ranges of seafloor 956 temperature, thermal gradient, source gas composition, 957 and gas venting rate. Our analysis shows that venting rate is 958 the principal control on vent gas and hydrate chemistry (Fig. 959 1). 960

The Bush Hill vent gas and hydrate chemistries are then 961 examined in terms of gas venting rates (Fig. 2). If venting 962 rates are slow ($q < \sim 0.5 \text{ kg/m}^2 \text{ a}$), hydrate crystallization 963 will not reach the sea floor. If venting rates are fast 964 $(q > \sim 20 \text{ kg/m}^2 \text{ a})$, the vent gas has almost the same 965 composition as the source gas, and the heaviest possible 966 hydrates are crystallized at the surface. Between these 967 extremes both the vent and hydrate gas compositions 968 depend strongly on venting rate. The hydrates sampled 969 from the Bush Hill mound are dryer (contain less $C_3 + C_4$) 970 than the hydrates that would crystallize from the vent gases 971 sampled over the last decade (Fig. 2). If the sampled 972 hydrates were crystallized more than 10 years ago, their 973 comparative dryness suggests that the gas flux was lower 974 (and the vent gas dryer) in the past. 975

Changes in vent gas chemistry observed at five repeat 976 submersible visits to the Bush Hill site are interpreted in 977 terms of venting rates (Fig. 4). The range in vent gas 978 chemistry is greater at any instant in time than the changes 979 in the average vent gas chemistry between submersible 980 visits. Nearly a two order of magnitude variation in venting 981 rate $(1-55 \text{ kg/m}^2 \text{ a})$ is suggested by the chemical differ-982 ences at any instant. We suggest that the different chemistry 983 of different bubble streams sampled at the same time reflects 984 different rates of movement across the HSZ. Changes in 985 average venting rate between submersible visits of a factor 986 of 2 over a few years are suggested by changes in the vent 987 chemistry. 988

Fig. 5 summarizes a suggestion of how venting systems 989 such as that at Bush Hill may in general evolve. In the initial 990 mud volcano stage of venting, total gas venting, Q, is 991 maximum but spread over a broad area. With time the gas 992 venting is increasingly organized into fewer venting of 993 smaller cross-sectional areas, and the gas flux (mass of gas 994 crossing a unit area per unit time), q, increases. This 995 increasing organization is reflected in the chemical disparity 996 between the vent gas sampled in the last decade and the 997 older hydrates. Eventually, the Bush Hill venting may 998 decrease to the point hydrates no longer crystallize at 999 the surface, and the hydrate mound will dissolve and be 1000 partially replaced by carbonate. 1001

There are many uncertainties in our modeling analysis 1002 and the suggestions we draw from it, and an important 1003 aspect of the vent-rate-control-on-vent-chemistry hypothesis is therefore that it can be tested by monitoring the 1005 chemistry and venting rate of a single gas stream, by further 1006 documenting variation in the venting rate of individual 1007 bubble streams at a single site during the same submersible 1008

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visit, or by drilling the Bush Hill (or a similar) hydrate 1009 mound to determine if the gas and hydrate chemical 1010 variability changes with depth as predicted in Fig. 3. 1011

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